ctools - Feature #2713

implementation of a time-variability search tool

11/05/2018 04:30 PM - Bonnefoy Simon

Status:	Closed	Start date:	11/05/2018	
Priority:	Normal	Due date:	11/11/2018	
Assigned To:	Bonnefoy Simon	% Done:	100%	
Category:		Estimated time:	0.00 hour	
Target version:	1.6.0			
Description				
Implementation of a new tool to search for time variability. The tool is intended to be used to search for variability over an event list. The energy range and the time scale can be directly set by the user. The tool is based on the algorithms used in HEGS and FAVA.				
See here for more details:				
nttps://cta-redmine.irap.omp.eu/attacnments/download/2963/variability_presentation.pdf				

History

#1 - 11/21/2018 11:13 AM - Bonnefoy Simon

- File skymap_three_sources.png added

- File skymap_variability.png added

- File variability_evolution.png added

The ctfindvar tool has been implemented.

The tool searches for variability over a given field of view. The user can provide some observation files or an event list. The position of several sources can be specified in an xml file, or only one source coordinate can be set as parameter. In order to search for variability, the tool bins the events in time for each pixel, and compares the number of events in each bin with respect to the others bins, set as background. Only bins with a significance lower than 4.5, with respect to the background sample, are kept for the estimation of the background. This significance threshold is a parameter of the tool and can be changed by the user. The energy boundaries and the time scale can be set by the user as parameters.

The tool produces one FITS file as output containing: -a skymap with the highest significance for each pixel -the significance evolution of the source(s) -the max significance in the field of view for each time bin.

As an example, we simulated three sources: -two sources with different level of variability -one source is not variable

Here is the skymap obtained running ctskymap on the input files (count map).



The three sources are visible with different intensities.

Below is the skymap representing the significance of the variability for each pixel

5:00.0 50:00.0 45:	00.0 40:00.0 35:0	0,05:30 00.025:0	0.0 20:00.0 15:00.0
0 0 0			
24:00			
0,00			
22:0			
0: 0: 0: 0:		1.19	
20:00			
0 <u></u>			
18:0			

Here, only the two variable sources are detected, at different level, given the distinct varibality set for the two sources.

For the sources set in a xml file, a list of the significance evolution is contained in the output file. From this list, a lightcurve can be built together with the significance distribution (see exmaples/show_variability_evolution.py).



#2 - 11/21/2018 03:42 PM - Bonnefoy Simon

- Status changed from New to Pull request
- % Done changed from 0 to 100

The final tool can be found here:

Pull repo: simonbonnefoy/ctools Pull branch: 2713-variability-tool

#3 - 12/11/2018 02:30 PM - Knödlseder Jürgen

- Status changed from Pull request to Feedback
- Target version set to 1.6.0

I merged the code into the devel branch.

I did a bit of restructuring of the code to make it more modular. Specifically:

- I changed ctfindvar::init_cube to ctfindvar::create_cube that also includes the filling of the cube and the optional smoothing
- I added ctfindvar::analyse_cube that contains the analysis code that previously was in the ctfindvar::run method
- I added ctfindvar::get_pixels that returns a vector of sky map pixel indices
- I added ctfindvar::sky_model that builds a sky model for a given sky direction
- I changed the interface to ctfindvar::get_variability_sig that now returns the results in a GNdarray
- I changed ctfindvar::fill_alpha_vector to ctfindvar::get_alphas that now returns a vector of alpha values
- I added ctfindvar::get_tstart and ctfindvar::get_tstop methods to determine start and stop time
- I merged ctfindvar::write_srchist and ctfindvar::write_srchist_fits into a single ctfindvar::write_source_histograms method. The format of the
 binary table extensions was also changed, since it was not very clear what the content of the second position extension was. Now, the third
 extension contains three columns, a first with a source name, and two others with Right Ascension and Declination. One could image to store
 also TSTART and TSTOP in seconds, and to have the time reference saved in the header file, which would make the extension more OGIP
 compliant.

The interface of the class now looks as follows:

* @class ctfindvar

* @brief Time variability search tool

* @todo Add tool description.

class ctfindvar : public ctobservation {

public:

// Constructors and destructors
ctfindvar(void);
explicit ctfindvar(const GObservations& obs);
ctfindvar(int argc, char *argv[]);
ctfindvar(const ctfindvar& app);
virtual ~ctfindvar(void);

// Operators
ctfindvar& operator=(const ctfindvar& app);

// Methods
void clear(void);
void run(void);
void save(void);

// Get the information on the time interval from int time2inx(const GTime& time); GGti inx2gti(const int& indx); const GSkyMap& counts(void);

protected:

// Protected	d methods
void	init_members(void);
void	copy_members(const ctfindvar& app);
void	free_members(void);

VOID	get_parameters(void);
void	init_gtis(void);
void	create_cube(void);
void	fill_cube(GCTAObservation* obs);
void	analyse_cube(void);
std::vector <in< td=""><td>t> get_pixels(void);</td></in<>	t> get_pixels(void);
GNdarray	get_variability_sig(const int& ipix);
std::vector <d< td=""><td>ouble> get_alphas(const int& ipix) const;</td></d<>	ouble> get_alphas(const int& ipix) const;
double	gti_overlap(const GGti& gti1, const GGti& gti2) const;
GTime	get_tstart(void);
GTime	get_tstop(void);
GModelSky	sky_model(const GSkyDir& dir) const;
void	write_source_histograms(GFits& fits);

// Protected members

GSkyMap	m_counts;	//!< Counts for each time interval
std::vector <g< td=""><td>iGti> m_gti;</td><td>//!< List of time intervals</td></g<>	iGti> m_gti;	//!< List of time intervals
GModels	m_inmodel;	//!< List of models for source positions
GSkyDir	m_max_sig_dir;	//!< Sky direction associated with maximum significance
double	m_minoff; //	/!< Minimum counts for use in significance calculation
double	m_sig_threshold;	//!< Minimum significance required to set source as variable
GSkyMap	m_peaksigma	b; //!< Skymap holding the maximum significance
GNdarray	m_pixsigsrc;	<pre>//!< Store distributions of the source significances</pre>
GNdarray	m_pixsigmax;	//!< Store distribution for pixel with max significance
GTime	m_tstart; //	!< Start time for variability study
GTime	m_tstop; /	/!< Stop time for variability study
GEnergy	m_emin;	//!< Minimum energy for events
GEnergy	m_emax;	//!< Maximum energy for events
GModels	m_model_abov	e_thr; //!< Model storing position with significance above thr

};

Furthermore, I changed some of the parameter names for compliance with the other ctools. Specifically:

• I removed the prefix parameter and replaced it by the explicit filename parameters outmap and outmodel

- I renamed sig_thr to threshold
- I renamed smoothkrnl to smooth kernel
- I renamed smoothpar to smooth rad

I addition, the emin and emax parameters can no longer be zero, since the energy limits are needed in the alpha computation. This should be changed in the future so that general background models can be supported, and not only the IRF background model. I created a new feature for that: #2769.

#

General parameters

- inobs, f, a, events.fits,,, "Input event list or observation definition XML file"
- inmodel, f, h, NONE,,, "Input model definition file for extracting source positions"
- outcube, f, h, NONE,,, "Output counts cube file"
- outmap, f, a, map.fits,,, "Output map file"
- outmodel, f, a, model.xml,,, "Output model definition file"
- caldb, s, a, prod2,,, "Calibration database"
- irf, s, a, South_0.5h,,, "Instrument response function"

#

Variability search parameters

ysrc, r, a, 22.01,-90,90, "Second coordinate of the source of interest (DEC or galactic b)"

emin, r, a, 0.1,,, "Lower energy limit (TeV)"

emax, r, a, 100.0,,, "Upper energy limit (TeV)"

threshold, r, h, 4.5,,, "Significance threshold for variability detection"

minoff, r, h, 0,,, "Minimum required events in off bin for computing Noff and alpha"

smooth_kernel, s, h, NONE, GAUSSIAN|DISK|NONE,, "Smoothing kernel for counts maps" smooth_rad, r, h, 0.05,,, "Smoothing kernel radius (deg)"

#

Time binning parameters

tinterval, r, a, 300,,, "Time interval for each bin (seconds)"

tmin, t, a, NONE,,, "Start time (UTC string, JD, MJD or MET in seconds)"

tmax, t, a, NONE,,, "Stop time (UTC string, JD, MJD or MET in seconds)"

#

Spatial binning parameters

#== usepnt, b, h, no,,, "Use pointing instead of xref/yref parameters?"

nxpix, i, a, 200,,, "Size of the X axis in pixels" nypix, i, a, 200,,, "Size of the Y axis in pixels"

binsz, r, a, 0.02,,, "Image scale (in degrees/pixel)"

proj, s, a, CAR,AIT|AZP|CAR|GLS|MER|MOL|SFL|SIN|STG|TAN,, "Projection method"

- xref, r, a, 83.63,0,360, "First coordinate of image center in degrees (RA or galactic I)"
- yref, r, a, 22.51,-90,90, "Second coordinate of image center in degrees (DEC or galactic b)"
- #
- # Standard parameters

- nthreads, i, h, 0,0,1000, "Number of parallel processes (0=use all available CPUs)"
- chatter, i, h, 2,0,4, "Chattiness of output"
- clobber, b, h, yes,,, "Overwrite existing output files with new output files?"
- debug, b, h, no,,, "Activate debugging mode?" mode, s, h, ql,,, "Mode of automatic parameters"
- logfile, f, h, ctfindvar.log,,, "Log filename"

From my side things look okay. Could you please double check that the code is also okay from your side.

#4 - 03/06/2019 01:56 PM - Knödlseder Jürgen

- Status changed from Feedback to Closed

Code was merged into devel.

Files

skymap_variability.png	56 KB	11/21/2018	Bonnefoy Simon
skymap_three_sources.png	74.2 KB	11/21/2018	Bonnefoy Simon
variability_evolution.png	189 KB	11/21/2018	Bonnefoy Simon