

## ctools - Action #3929

### Follow up Masha's problems with the light curve of PWN2059

12/10/2021 03:15 PM - Knödlseider Jürgen

<b>Status:</b>	Closed	<b>Start date:</b>	12/10/2021
<b>Priority:</b>	Normal	<b>Due date:</b>	
<b>Assigned To:</b>	Knödlseider Jürgen	<b>% Done:</b>	100%
<b>Category:</b>		<b>Estimated time:</b>	0.00 hour
<b>Target version:</b>	2.0.0		
<b>Description</b>			
This action refers to an e-mail exchange with Masha concerning PWN2059 in the GPS simulation data.			

#### History

##### #1 - 12/10/2021 11:00 PM - Knödlseider Jürgen

- Status changed from New to In Progress

- % Done changed from 0 to 10

First I tried to reproduce Masha's results and I managed to do so:

```
2021-12-10T16:11:40: +=====+
2021-12-10T16:11:40: | Maximum likelihood optimisation results |
2021-12-10T16:11:40: +=====+
2021-12-10T16:11:40: === GOptimizerLM ===
2021-12-10T16:11:40: Optimized function value ..: 348541.336
2021-12-10T16:11:40: Absolute precision ..: 0.005
2021-12-10T16:11:40: Acceptable value decrease ..: 2
2021-12-10T16:11:40: Optimization status ..: converged
2021-12-10T16:11:40: Number of parameters ..: 231
2021-12-10T16:11:40: Number of free parameters ..: 4
2021-12-10T16:11:40: Number of iterations ..: 1
2021-12-10T16:11:40: Lambda ..: 0.0001
2021-12-10T16:11:40: Total number of iterations : 8
2021-12-10T16:11:40: Maximum log likelihood ..: -348541.336
2021-12-10T16:11:40: Observed events (Nobs) ..: 52107.000
2021-12-10T16:11:40: Predicted events (Npred) ..: 51863.589 (Nobs - Npred = 243.411000453176)
...
2021-12-10T16:11:40: === GModelSky ===
2021-12-10T16:11:40: Name ..: R000085_00001_J1737-3837
2021-12-10T16:11:40: Instruments ..: all
2021-12-10T16:11:40: Test Statistic ..: 34.7896858851891
2021-12-10T16:11:40: Observation identifiers ..: all
2021-12-10T16:11:40: Model type ..: PointSource
2021-12-10T16:11:40: Model components ..: "PointSource" * "ExponentialCutoffPowerLaw" * "Constant"
2021-12-10T16:11:40: Number of parameters ..: 7
2021-12-10T16:11:40: Number of spatial par's ..: 2
2021-12-10T16:11:40: RA ..: 264.370455982881 deg (fixed,scale=1)
2021-12-10T16:11:40: DEC ..: -38.6193821817504 deg (fixed,scale=1)
2021-12-10T16:11:40: Number of spectral par's ..: 4
2021-12-10T16:11:40: Prefactor ..: 9.37556142183426e-19 +/- 2.55043578929073e-19 [0,infty] ph/cm2/s/MeV (free,scale=1e-18,gradient)
2021-12-10T16:11:40: Index ..: -1.93913156282528 [-9,0] (fixed,scale=-1,gradient)
2021-12-10T16:11:40: CutoffEnergy ..: 2194778.42864015 [10000,500000000] MeV (fixed,scale=1000000,gradient)
2021-12-10T16:11:40: PivotEnergy ..: 1000000 MeV (fixed,scale=1000000,gradient)
2021-12-10T16:11:40: Number of temporal par's ..: 1
2021-12-10T16:11:40: Normalization ..: 1 (relative value) (fixed,scale=1,gradient)
2021-12-10T16:11:40: Number of scale par's ..: 0
```

**#2 - 12/10/2021 11:12 PM - Knödseder Jürgen**

- % Done changed from 10 to 20

Then I tried fitting the data of the first light curve bin with different combinations of model components. In all cases, the predicted number of counts correspond to the observed one. Even with only a few components, PWN 2059 got a low flux, comparable to Masha's run.

Components	Nobs-Npred	PWN 2059
Background only	4.69e-06	-
Background & IEM	0.0037	-
Background, IEM & PWN 2059	0.0059	9.51e-19 ± 2.57e-19
Background, IEM & PWN 2059 (free index)	0.0034	9.30e-19 ± 2.60e-19

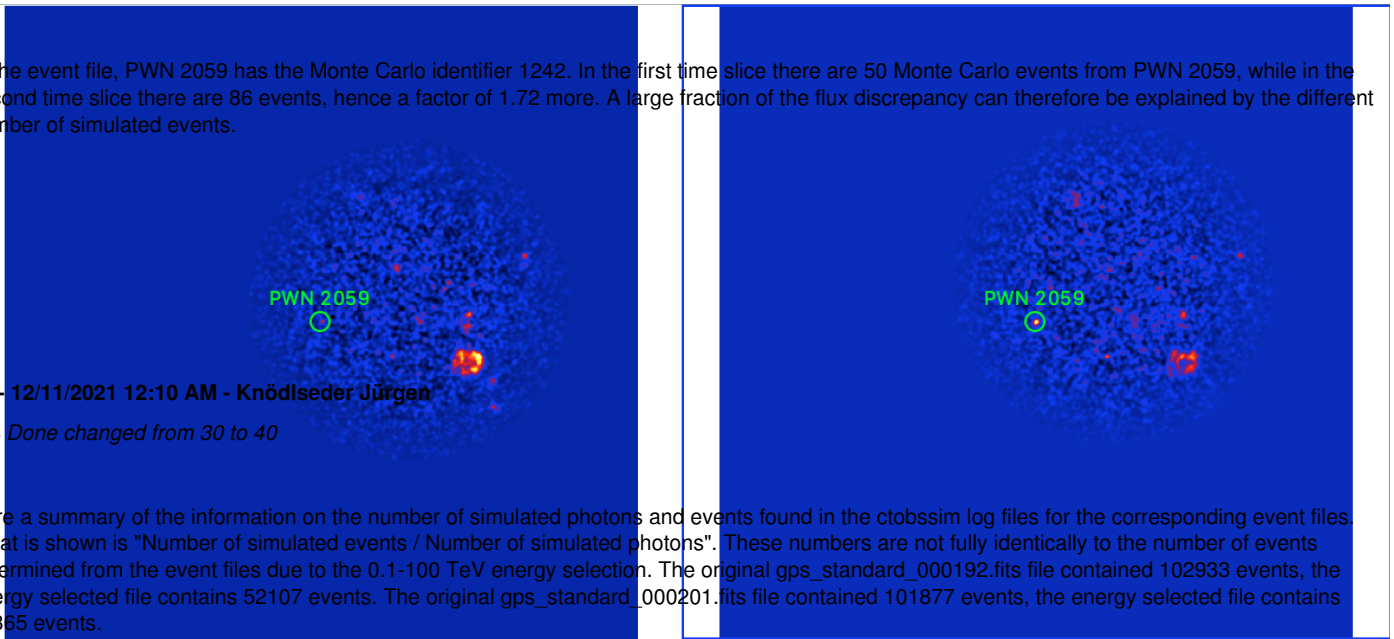
**#3 - 12/10/2021 11:45 PM - Knödseder Jürgen**

- File pwn2059\_192\_201\_smoothed.png added

- % Done changed from 20 to 30

Here are smoothed background subtracted sky maps of the two time bins 192 and 201 (files gps\_standard\_000192.fits and gps\_standard\_000201.fits). Obviously PWN 2059 appears much brighter in the second time bin compared to the first time bin.

In the event file, PWN 2059 has the Monte Carlo identifier 1242. In the first time slice there are 50 Monte Carlo events from PWN 2059, while in the second time slice there are 86 events, hence a factor of 1.72 more. A large fraction of the flux discrepancy can therefore be explained by the different number of simulated events.



**#4 - 12/11/2021 12:10 AM - Knödseder Jürgen**

- % Done changed from 30 to 40

Here a summary of the information on the number of simulated photons and events found in the ctobssim log files for the corresponding event files. What is shown is "Number of simulated events / Number of simulated photons". These numbers are not fully identical to the number of events determined from the event files due to the 0.1-100 TeV energy selection. The original gps\_standard\_000192.fits file contained 102933 events, the energy selected file contains 52107 events. The original gps\_standard\_000201.fits file contained 101877 events, the energy selected file contains 51365 events.

In the 3rd to 6th energy slices, the number of simulated events is systematically larger in the 201 time bin compared to the 192 time bin. Maybe this is just "bad luck"?

Energy range	192	201
30 - 72.33 GeV	0/778	0/842
72.33 - 174.37 GeV	20/751	17/681
174.37 - 420.38 GeV	17/441	34/441
420.38 GeV - 1.01 TeV	10/311	15/317
1.01 - 2.44 TeV	4/158	13/187
2.44 - 5.89 TeV	1/35	8/35

5.89 - 14.20 TeV	1/4	0/2
14.20 - 34.24 TeV	0/0	0/0
34.24 - 82.54 TeV	0/0	0/0
82.54 - 199 TeV	0/0	0/0
Total	53/2478	87/2505

#### #5 - 12/11/2021 12:45 AM - Knödseder Jürgen

I tried to reproduce the simulation on my Mac using the ctools development version with the aim to study simulation procedure. Apparently the use of the random number generator changed, hence the code did not reproduce exactly the same number of events. Switching back to version 1.7.0 that was used for the GPS simulations I managed to reproduce the number of photons and events that were simulated for PWN 2059 in the GPS run 192. I therefore will use ctools version 1.7.0 for the following studies.

Simulation	Seed	Photons	Events
Original GPS dataset (ctools-1.7.0)	192	2478	53
Mac OS X (ctools-2.0.0.dev)	192	2500	50
Mac OS X (ctools-1.7.0)	192	2478	53
Kepler (ctools-1.7.0)	192	2478	53

#### #6 - 12/13/2021 11:10 AM - Knödseder Jürgen

- File *sim192.png* added
- File *pwn2059\_reference.png* added
- % Done changed from 40 to 50

{UPDATE} The two leftmost plots are so far only for 324 simulations. I will try to make 1000 simulations to verify the match with the expectation, yet already with 324 simulations there seems to be a reasonable match, and no significant difference to the faster simulations where only PWN 2059 is used.

I investigated whether the observed low number of events is explained by pure statistical fluctuations, or whether there is a possible issue with the random number generator.

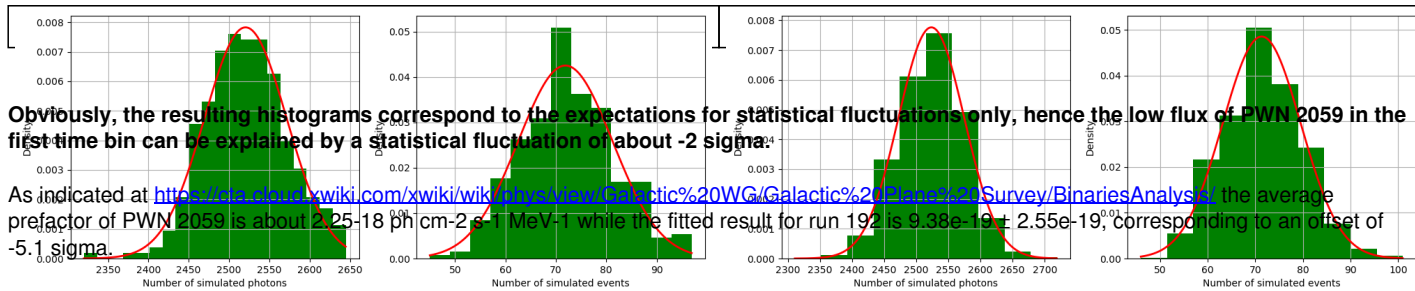
To do this I run the simulations using ctools version 1.7.0 by changing the seed parameter of ctobssim which sets the seed value of the random number generator, and extract the number of photons and events that are simulated for PWN 2059. Repeating the process 1000 times for different seed values gives then a distribution of number of photons and events that can be compared to the expectation based on statistical fluctuations only.

The outcome of this process are show below, with the left two panels corresponding to the simulation of the full GPS model, reproducing exactly the simulation that was done for the GPS, while the right two panels correspond to simulations of only PWN 2059, speeding up considerably the simulations. The table below gives the moments of the photon and event distributions. The observed number of 53 events for PWN 2059 in run 192 is a deviation of -18.89 events from the mean, corresponding to -2.01 standard deviations.

Parameter	Value
Mean number of photons	2520.50
Standard deviation in number of photons	50.97
Mean number of events	71.89

Standard deviation in number of events

9.38



Obviously, the resulting histograms correspond to the expectations for statistical fluctuations only, hence the low flux of PWN 2059 in the first time bin can be explained by a statistical fluctuation of about -2 sigma.

As indicated at <https://ins.cfa.harvard.edu/wiki/wiki/phys/view/Galactic%20WG/Galactic%20Plan%20Survey/BinariesAnalysis> the average prefactor of PWN 2059 is about  $2.25 \cdot 10^{-18} \text{ ph cm}^{-2} \text{ s}^{-1} \text{ MeV}^{-1}$  while the fitted result for run 192 is  $9.38 \cdot 10^{-19} \pm 2.55 \cdot 10^{-19}$ , corresponding to an offset of -5.1 sigma.

So the question is why a statistical fluctuation of about -2 sigma translates into a prefactor offset of -5.1 sigma?

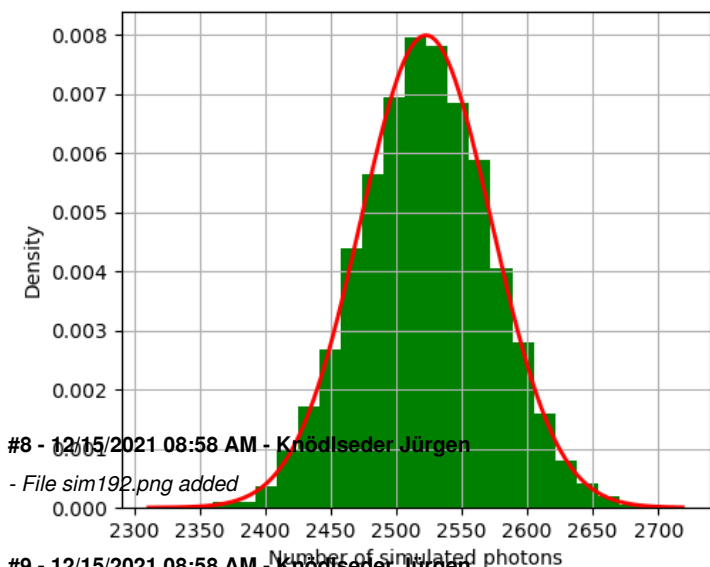
Interestingly, 53 counts is only a factor of 0.74 lower than 72 counts, in other words, the mean prefactor expected for run 192 would be  $9.38 \cdot 10^{-19} / 0.74 = 1.27 \cdot 10^{-18}$  which is almost a factor of 2 below the mean value found by Masha for the other bins.

#7 - 12/15/2021 08:53 AM - Knödseder Jürgen

- File pwn2059\_10000sims.png added

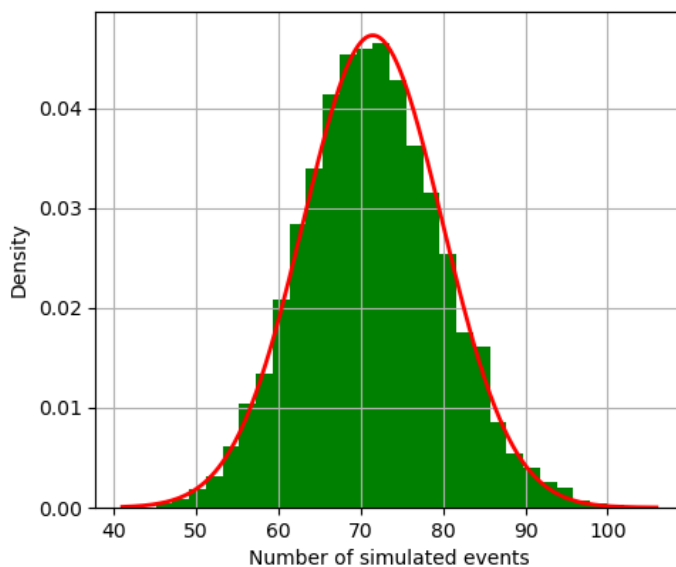
Below is the result for 10000 simulations of PWN 2059 only. The moments of the distributions are given in the Table below. The observed number of 53 events for PWN 2059 in run 192 is a deviation of -18.44 events from the mean, corresponding to -2.18 standard deviations.

Parameter	Value
Mean number of photons	2522.72
Standard deviation in number of photons	49.91
Mean number of events	71.44
Standard deviation in number of events	8.44



#8 - 12/15/2021 08:58 AM - Knödseder Jürgen

- File sim192.png added



#9 - 12/15/2021 08:58 AM - Knödseder Jürgen

- File deleted (sim192.png)

#10 - 12/15/2021 09:01 AM - Knödlseeder Jürgen

- File *sim192.png* added

#11 - 12/15/2021 09:01 AM - Knödlseeder Jürgen

- File deleted (*sim192.png*)

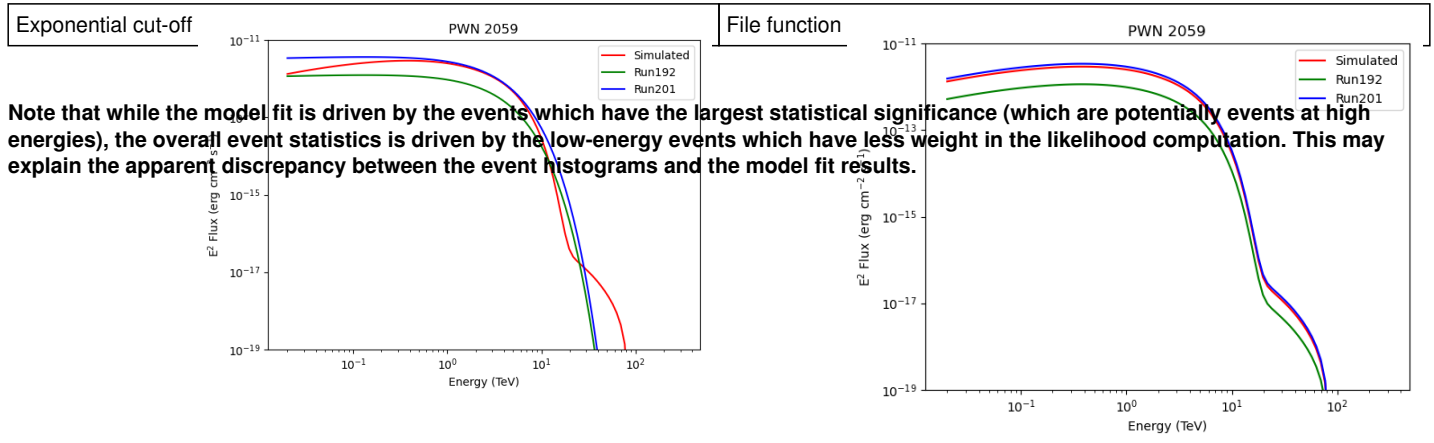
#12 - 12/15/2021 10:28 AM - Knödlseeder Jürgen

- File *sed\_cat\_model.png* added

- File *sed\_mc\_model.png* added

Below the results of fitting various models to the runs 192 and 201. For illustration, the fitted SEDs for the exponential cut-off model (left) and of the MC file function model (right) are shown in the two panels below.

Model	XML	TS (Run192)	Norm (Run192)	TS (Run201)	Norm (Run201)
Catalogue model (exp. cutoff)	mc_model_ready_roi_width.xml	34.79	$0.94 \pm 0.26$	209.76	$2.75 \pm 0.36$
MC model (file function)	model_cat_pwn2059.xml	34.55	$0.39 \pm 0.11$	209.28	$1.16 \pm 0.15$
Idem, free bkg index and IEM	model_cat_pwn2059_free-index-iem.xml	35.02	$0.40 \pm 0.11$	209.58	$1.17 \pm 0.15$



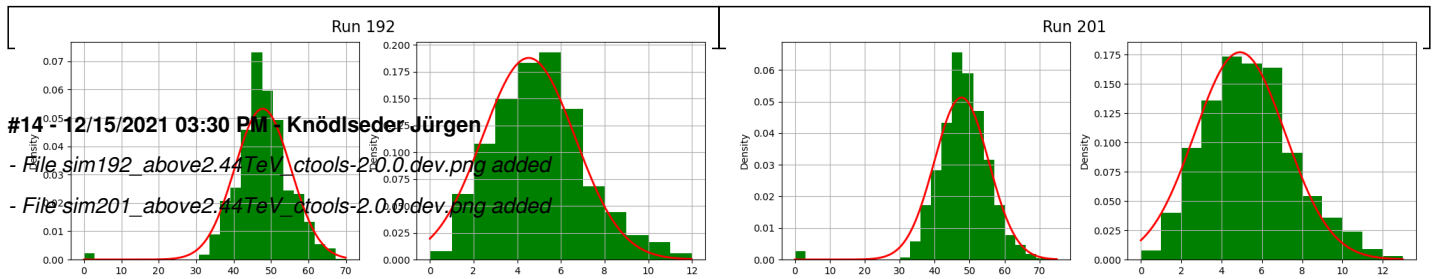
#13 - 12/15/2021 01:24 PM - Knödlseeder Jürgen

- File *sim192\_above2.44TeV.png* added

- File *sim201\_above2.44TeV.png* added

I now investigated the distribution of events with changing random number generator seed for energies above 2.44 TeV, where in run 192 2 events are simulated while in run 201 8 events are simulated. The distribution for 1000 simulations are shown in the plots below, moments of the distribution are given in the following table. While 2 events are on the left tail of the distribution, 8 events are on the right tail of the distribution. Note that for some reason, 8 out of 1000 seed values produce zero photons (seed values 81, 192, 222, 294, 465, 557, 703, and 888).

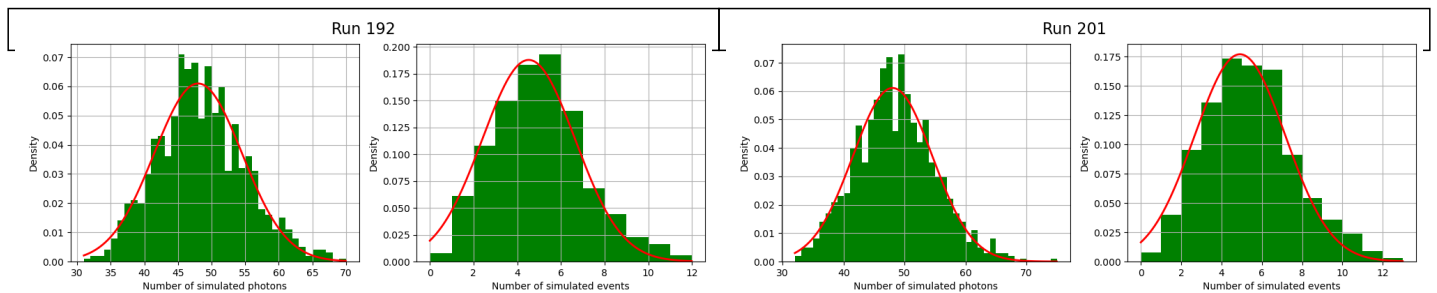
Parameter	Run 192 value	Run 201 value
Mean number of photons	47.77	47.69
Standard deviation in number of photons	7.49	7.77
Mean number of events	4.53	4.92
Standard deviation in number of events	2.13	2.26



#14 - 12/15/2021 03:30 PM - Knödlseher Jürgen  
 - File sim192\_above2.44TeV\_ctools-2.0.0.dev.png added  
 - File sim201\_above2.44TeV\_ctools-2.0.0.dev.png added

Repeating the simulations for ctools-2.0.0.dev does not show the zero photon bins. The mean photon number is slightly larger, its standard deviation is slightly reduced. Not clear what the difference between ctools-1.7.0 and ctools-2.0.0.dev regarding simulations is.

Parameter	Run 192 value	Run 201 value
Mean number of photons	48.00	48.03
Standard deviation in number of photons	6.54	6.53
Mean number of events	4.53	4.92
Standard deviation in number of events	2.13	2.26



#### #15 - 12/15/2021 03:45 PM - Knödlseher Jürgen

Below a table with the source significance in a given time bin, computed from the TS value ( $\sqrt{\text{TS}}$ ), the fitted normalisation ( $\text{norm}/\text{error}$ ) and the number of source counts ( $\sqrt{\text{events}}$ ). It is a bit surprising that the first two values differ so much, I would have expected that to first order both values are similar. It is also surprising that the TS significance for run 201 is larger than the event significance, since the event significance does not account for background events, hence it formally a upper limit to the significance of the source.

Run	Sigma(TS)	Sigma(Norm)	Sigma(events)
192	5.9	3.6	7.3
201	14.5	7.6	9.3

#### #16 - 12/15/2021 04:33 PM - Knödlseher Jürgen

Maybe the problem lies in the fitted flux for PWN 2059.

Below a table specifying for the first two runs the simulated number of events and the predicted number of events from the best fitting file function model. **For some reason the best fitting model value for run 192 results in a predicted number of events that is about a factor of two below the level that is expected from the simulated number of events.** No significant discrepancy is observed for run 201. As a reminder, the fitted normalisation of the file function is given in parenthesis with the associated statistical error.

I also did the analysis with energy dispersion enabled, which did not change the results.

Hence there are actually two effects that contribute to the low flux point for run 192:

- first the number of simulated counts for run 192 is on the low side, yet the number is compliant with expectations from statistical fluctuations
- second the fitted flux is about a factor two too low, underpredicting the number of events for PWN 2059. **It is this second effect that still needs to be understood.**

Run	Nsim	Npred	Npred (edisp)
192	53	26.250 (0.39 ± 0.11)	26.751 (0.39 ± 0.11)
201	87	85.522 (1.16 ± 0.15)	87.089 (1.16 ± 0.15)

#### #17 - 12/16/2021 04:34 PM - Knödseder Jürgen

I now fitted the data of run 192 in the same 10 energy bins that are used internally by ctobssim for the simulations in order to investigate whether the fitted normalisation factor changes with energy. The results are summarised in the table below. Between 0.07 and 1 TeV the number of predicted events is significantly lower than the number of observed events, only at larger energies both quantities become comparable. The discrepancy increases globally with decreasing energy. I also did the analysis with energy dispersion enabled, yet this did not significantly change the results.

Emin	Emax	Norm	Nsim	Npred		Norm (edisp)	Npred (edisp)
0.03	0.07	1.00	0	0.00		0	0.00
0.07	0.17	0.46 ± 0.59	20	5.10		0.42 ± 0.54	5.15
0.17	0.42	0.39 ± 0.23	17	8.66		0.39 ± 0.23	8.70
0.42	1.01	0.34 ± 0.17	10	6.29		0.34 ± 0.17	6.29
1.01	2.44	0.45 ± 0.22	4	4.90		0.45 ± 0.23	4.91
2.44	5.89	0.18 ± 0.25	1	0.74		0.18 ± 0.25	0.74
5.89	14.20	3.32 ± 3.37	1	0.99		3.22 ± 3.26	0.99
14.20	34.24	1.00	0	0.00		1.00	0.00
34.24	82.54	1.00	0	0.00		1.00	0.00
82.54	199.00	1.00	0	0.00		1.00	0.00

I then replaced the radial disk source by a point source to check whether the issue is related to the use of a small extended source. The results are basically identical, excluding the radial disk model as the origin of the discrepancy.

Emin	Emax	Norm	Nsim	Npred
0.03	0.07	1.00	0	0.00
0.07	0.17	0.46 ± 0.59	20	5.10
0.17	0.42	0.39 ± 0.23	17	8.64
0.42	1.01	0.34 ± 0.17	10	6.28
1.01	2.44	0.45 ± 0.22	4	4.88
2.44	5.89	0.18 ± 0.25	1	0.74
5.89	14.20	3.32 ± 3.37	1	0.99
14.20	34.24	1.00	0	0.00
34.24	82.54	1.00	0	0.00
82.54	199.00	1.00	0	0.00

I also analysed run 201 in the same way to check the results for a "normal" run. Except for the 0.07-0.17 TeV bin, where the number of predicted events is underestimated, the number of predicted events is comparable to the number of simulated events.

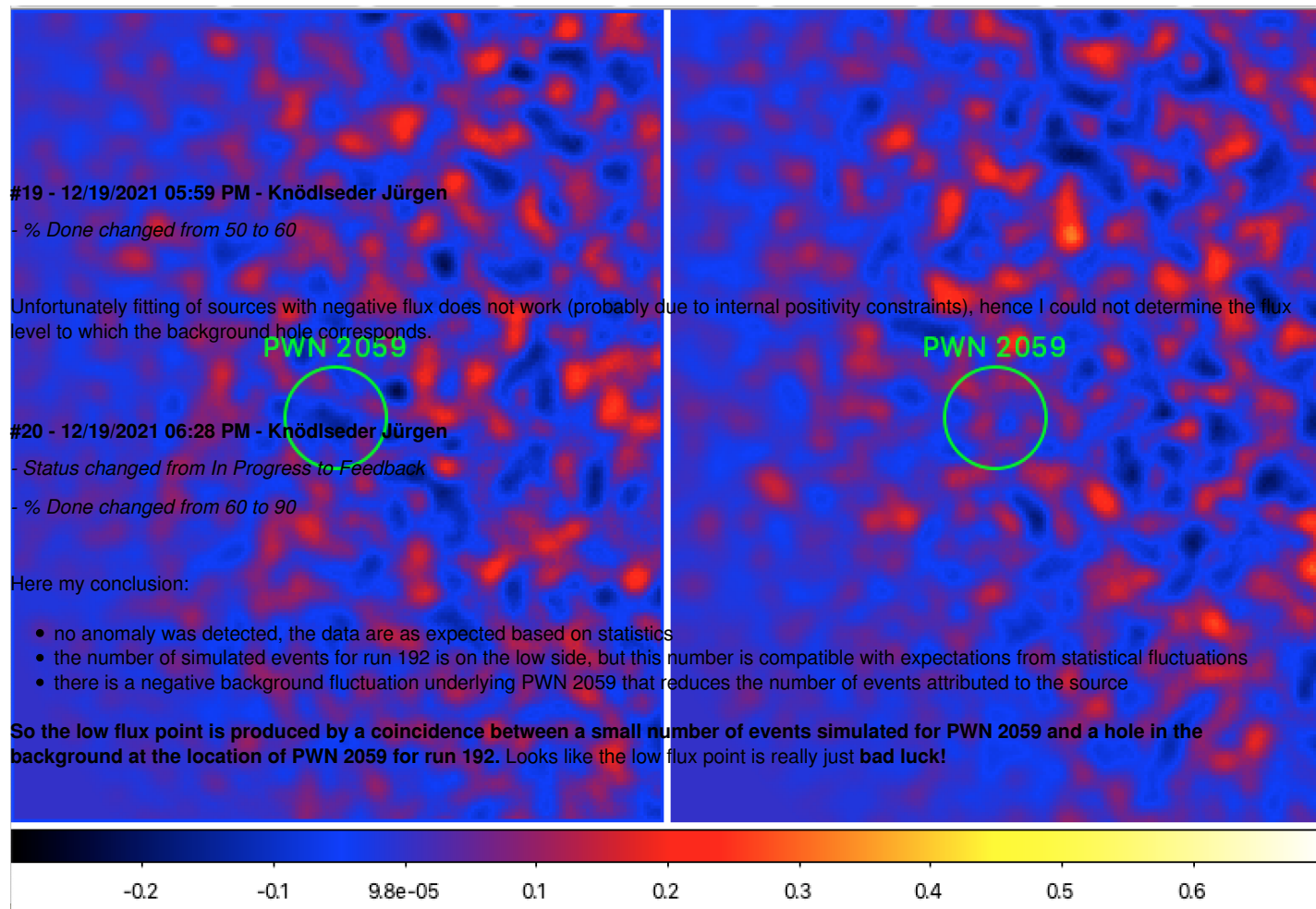
Emin	Emax	Norm	Nsim	Npred		Norm (edisp)	Npred (edisp)
0.03	0.07	1.00	0	0.00		0	0.00
0.07	0.17	0.66 ± 0.59	17	8.58		0.59 ± 0.54	8.34
0.17	0.42	1.57 ± 0.32	34	38.83		1.57 ± 0.32	39.19
0.42	1.01	0.79 ± 0.22	15	15.59		0.78 ± 0.22	15.60
1.01	2.44	1.16 ± 0.33	13	13.47		1.16 ± 0.33	13.47
2.44	5.89	1.77 ± 0.65	8	7.73		1.77 ± 0.65	7.73
5.89	14.20	1.00	0	0.31		1.00	0.32
14.20	34.24	0.00	0	0.00		0.00	0.00

34.24	82.54	1.00 0.00	0	0.00		1.00	0.00
82.54	199.00	1.00	0	0.00		1.00	0.00

### #18 - 12/19/2021 05:48 PM - Knödlseeder Jürgen

- File skymap-nopwn.png added

I now removed the PWN 2059 events from the simulated events for run 192 and 201 using ctselect with the parameter `expr="MC_ID!=1242"`. This reveals the background underlying the source events. Sky maps using the selected events are shown in the panels below, left is for run 192, right is for run 201. Apparently, there is a lack of background events for run 192 in the region of PWN 2059 which may explain why the fitted flux is so low.



### #21 - 03/15/2022 12:48 PM - Knödlseeder Jürgen

- Status changed from Feedback to Closed

- % Done changed from 90 to 100

#### Files

File Name	Size	Date	Author
pwn2059_192_201_smoothed.png	353 KB	12/10/2021	Knödlseeder Jürgen
pwn2059_reference.png	44.5 KB	12/13/2021	Knödlseeder Jürgen
pwn2059_10000sims.png	43.9 KB	12/15/2021	Knödlseeder Jürgen
sim192.png	44.2 KB	12/15/2021	Knödlseeder Jürgen
sed_mc_model.png	33.5 KB	12/15/2021	Knödlseeder Jürgen
sed_cat_model.png	32.4 KB	12/15/2021	Knödlseeder Jürgen
sim192_above2.44TeV.png	42.8 KB	12/15/2021	Knödlseeder Jürgen
sim201_above2.44TeV.png	41.8 KB	12/15/2021	Knödlseeder Jürgen



sim192_above2.44TeV_ctools-2.0.0.dev.png	45.8 KB	12/15/2021	Knödseder Jürgen
sim201_above2.44TeV_ctools-2.0.0.dev.png	42.7 KB	12/15/2021	Knödseder Jürgen
skymap-nopwn.png	321 KB	12/19/2021	Knödseder Jürgen